### VLBI part II: Accessing and calibrating data



**EUROPEAN ARC** 

**ALMA Regional Centre** 

Design: Katharina Immer and Megan Lewis

Interactive Training in Reduction and Analysis of INterferometric data with the European ARC network



### I-TRAIN with the European ARC Network

# 24: VLBI Part II: Obtaining and calibrating VLBI data with phased ALMA

Tutors: Michael Janssen, Georgios-Filippos Paraschos, Kazi Rygl









### I-TRAIN on VLBI with ALMA

Introduction to VLBI: Science and Proposals (Part 1)

- mm VLBI science with ALMA
- proposal opportunities
- planning an ALMA VLBI proposal
- ALMA OT for a VLBI proposal

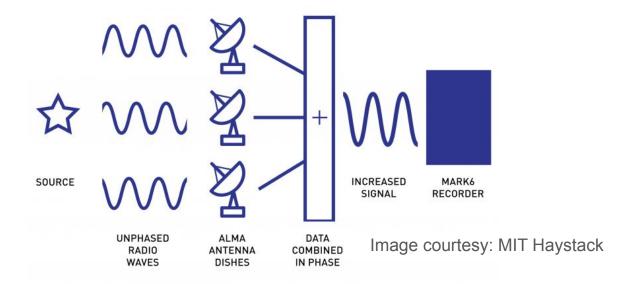
Working with ALMA VLBI data: Accessing and calibrating data (this talk)

- Phased ALMA data and what is in the ALMA Science Archive
- Obtaining GMVA
- Obtaining EHT data
- Calibrating GMVA and EHT data

# Phased ALMA participates to VLBI networks as large single dish thanks to the ALMA Phasing System



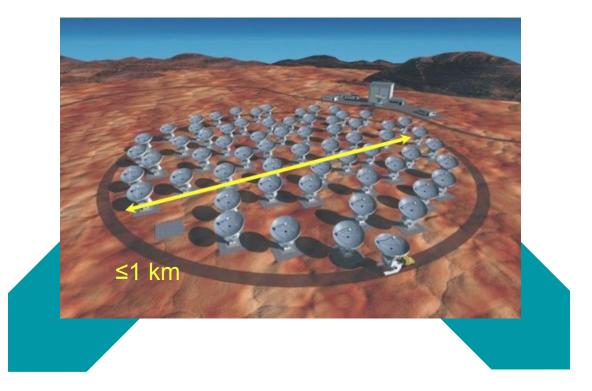
# **ALMA Phasing system**



APS computes the phase adjustments relative to a reference antenna. The phased signals are then summed in the ALMA correlator to create the (virtual) summed antenna.

See Matthews et al. 2018 for more details

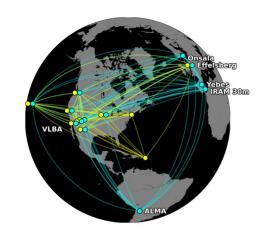
### Data collected during phased ALMA observations



~77m single dish to be correlated with VLBI antennas

ALMA 12-m Array interferometer

# Quality assessment of phased ALMA observations



ALMA: linear polarisation (X,Y)
VLBI antennas: circular polarisation (L,R)

$$V_{+\odot} = egin{pmatrix} V_{XR} & V_{XL} \ V_{YR} & V_{YL} \end{pmatrix} egin{pmatrix} {\sf QA2} \ {\sf PolConvert} \end{pmatrix} V_{\odot\odot} = egin{pmatrix} V_{RR} & V_{RL} \ V_{LR} & V_{LL} \end{pmatrix}$$

See Goddi et al. 2019, Marti-Vidal 2016 for more details

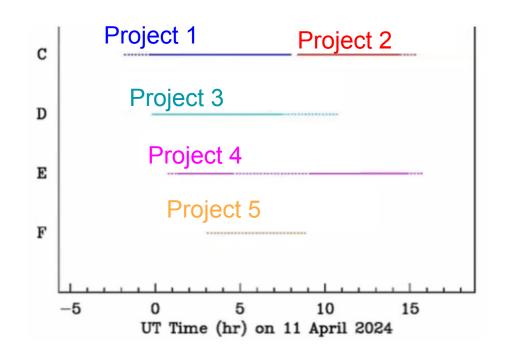
### Quality assessment of phased ALMA observations

#### Goals:

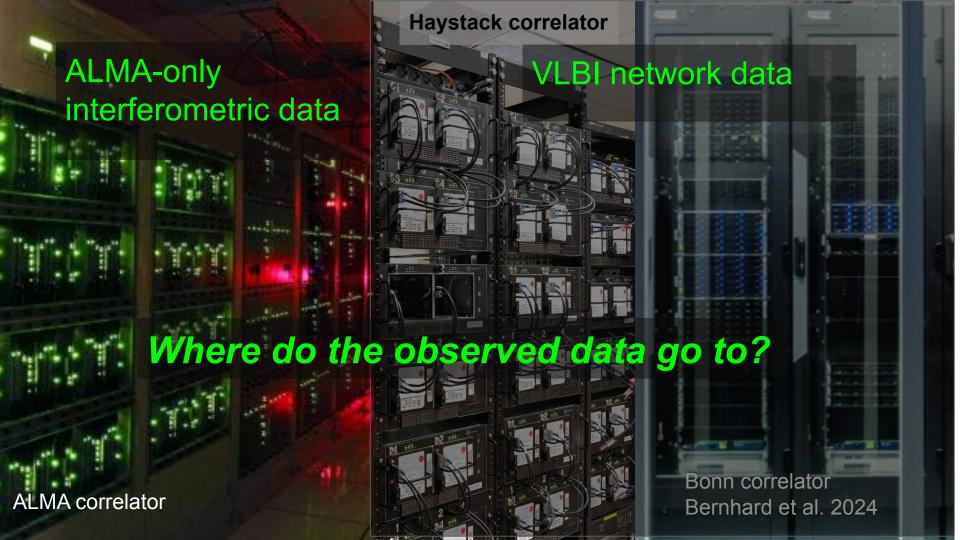
- polarisation conversion of mixed XY/LR visibilities to circular only to allow VLBI data calibration
- deliver ALMA interferometer data to ALMA Science Archive (project based)

# QA2 for ALMA VLBI data entails a regular polarisation calibration with a few exceptions:

- done track-based rather than project based;
   calibrators can be shared
- full details in Goddi et al. 2019







### **ALMA Science Archive**

Searching on project code "20\*.V" gives all VLBI projects (20xx.1.xxxxx.V)

© Observations (128)  Projects (5077)				
Project code: 20*.V ×	de ALMA source name	RA h:m:s - Dec d:m:s - Band	Cont.sens. mJy/beam Frequency support	↑ Release date Publication
□ ↔ 20*.V	× ×	ILIII.S DOC U.III.S DUIG	Contraction in System in requestery support	Therease date in ablication
○ ↔ ~ ₺ ፟ 2016.1.0	01290.V ngc1052	02:41:04.799 -08:15:20.7526	0.0214 212.181230.05 GH	Hz 2018-10-20 4
□ ↔ ↔ ∼ ₺ ፟ 2016.1.	01116.V OJ287	08:54:48.875 +20:06:30.64 3	0.0130 85.344101.213 GH	Hz 2018-10-20 4
$\bigcirc \Leftrightarrow \leftrightarrow \sim \not \bot \boxtimes \qquad \boxed{2016.1.6}$	01176.V 3C279	12:56:11.167 -05:47:21.528 6	0.0179 212.164230.031 G	Hz 2018-10-20 4
$\bigcirc \Leftrightarrow \leftrightarrow \sim \not \bot \boxtimes \qquad \boxed{2016.1.0}$	00413.V Sagittarius_A_star	17:45:40.032 -29:00:28.26 3	0.0114 85.323101.189 GH	Hz 2018-10-20 9
$\bigcirc \Leftrightarrow \leftrightarrow \sim \checkmark \boxtimes \bigcirc \bigcirc$	01198.V Centaurus_A	13:25:27.615 -43:01:08.80! 6	0.0133 212.163230.03 GH	Hz 2018-10-20 4
$\bigcirc \Leftrightarrow \leftrightarrow \sim \checkmark \boxtimes \qquad \boxed{2016.1.}$	01114.V OJ287	08:54:48.875 +20:06:30.64 6	0.0240 212.191230.061 G	Hz 2018-10-20 3
$\bigcirc \Leftrightarrow \leftrightarrow \sim \checkmark \boxtimes \qquad \boxed{2016.1.6}$	01176.V 3C279	12:56:11.167 -05:47:21.528 6	0.0162 212.166230.034 G	GHz 2018-10-20 4
$\bigcirc \leftrightarrow \sim \checkmark \boxtimes \bigcirc \bigcirc$	01176.V 3C279	12:56:11.167 -05:47:21.528 6	0.0191 212.166230.033 G	6Hz 2018-10-20 4
$\bigcirc \leftrightarrow \sim \checkmark \boxed{2016.1.0}$	01290.V ngc1052	02:41:04.799 -08:15:20.7526	0.0154 212.181230.05 GH	Hz 2018-10-20 4
$\bigcirc \Leftrightarrow \leftrightarrow \sim \bot$ $\bigcirc$ 2016.1.0	01404.V Sagittarius_A_star	17:45:40.036 -29:00:28.17(6	0.0133 212.139230.005 G	9Hz 2018-10-20 8

# ALMA-only VLBI data in the archive

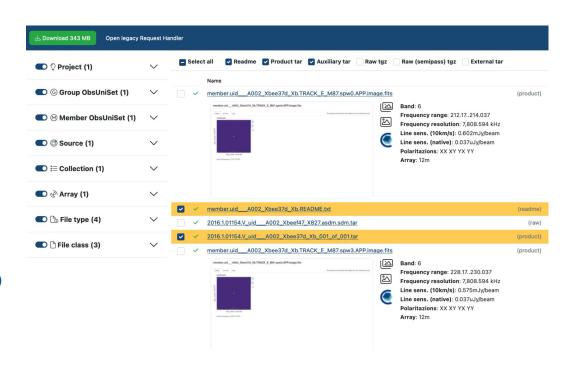
#### Raw data

Calibration script and the calibration tables from QA2

#### Image products:

- MFS full Stokes images, one per spectral window
- No cubes, no combined continuum image
- No calibrator images (these can be the target of another project in the same track)

Public 1 year after confirmed successful correlation at the VLBI correlator



#### VLBI AGN POLarization data with ALMA

Welcome to VAPOLA — the first ALMA AGN sources archive observed in APS mode, containing multi-epoch multi-band full Stokes data products ready for science.

#### **Introduction to VAPOLA project**

VAPOLA is a user friendly repository of ALMA data observed in APP mode. It is located in the Italian Astronomical Archives Center (INAF) servers (Trieste, Italy). It contains, ready-for-science heigh order products and to the whole community. Every year it will be updated with observations of the current and new sources observed in the different bands.

#### What to find in this webpage repository?

In this wabepage you will find an interactive portal to the data. Left side menu (hamburger icon in top left margin) contains the redirection for the Home (actual page), Download (where you will have access to the data), Documentation (paper, technical memo) and my contact email for any kind of question or problems you may have.

#### How to download the data

To download the data, you need to click the menu, and select the Download page. You will have access to the tree data structure. Data is organized as follows:

- 1. in the root folder, there are three folders VLBI, non-VLBI, documentation
- VLBI contains the data per Source, Year, Band, Day\_Month
- 3. non-VLBI contains the (generally short) observations of ALMA in non APP mode DURING the Global Campaings
- 4. documentaiton. Folder that contains csv files that summariz the data

You can download the data in three ways: individually, by wishing list (selecting one or more files), and filtering by different criteria ("Download by Selection")

#### Note for PI's of private data that have kindly shared with us

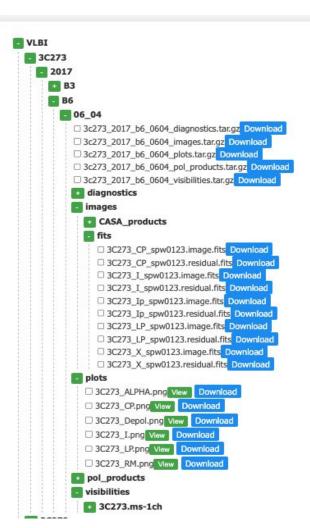
To access to protected data, you need to login in the Download page.

#### Contact

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webpage: <a href="http://vapola.ia2.inaf.it/">http://vapola.ia2.inaf.it/</a>

Piano Nazionale di Ripresa e Resilienza (PNRR) – Missione 4 "Istruzione e Ricerca" – Componente 2 "Dalla Ricerca all'Impresa" Investimento 1.1 "Fondo per il Programma Nazionale di Ricerca e Progetti di Rilevante Interesse Nazionale (PRIN)". PI: Ciriaco Goddi



### VAPOLA repository

Organised per source, per band, per observing run

Advanced polarisation data products (FITS)

Plots (PNG) of advanced data products plus spectral index and RM maps.

Calibrated visibilities (without channel smoothing)

### Example: M87 in polarisation

-10

Relative RA (arcsec)

12

8-

6

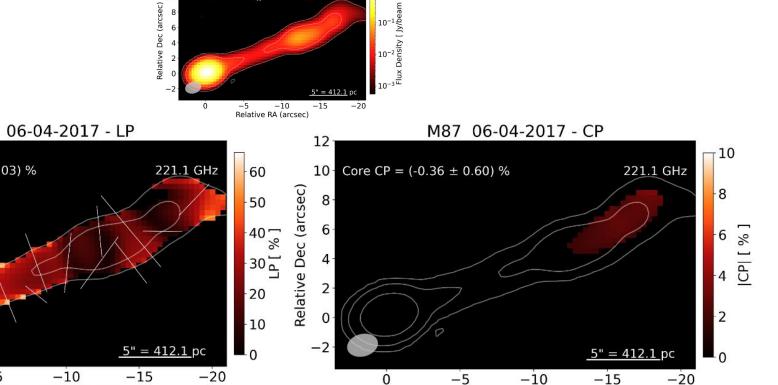
2-

Relative Dec (arcsec)

Core LP =  $(2.15 \pm 0.03)$  %

10 - Core I =  $(1.29 \pm 0.13)$  Jy/beam

221.1 GHz



# Example: M87 RM and spectral index

